

# Passage of particles through matter

Alexander Khanov

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Oklahoma State University

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# Nuclear interaction length

- As we've seen, high energy electrons and photons produce electromagnetic cascades, or “showers”
- Similarly, high energy hadrons (pions, kaons, protons, neutrons) produce “hadronic showers”
  - ▶ when hadrons hit nuclei, they create other particles due to strong interaction
- This process can be characterized by a typical length parameter called “nuclear interaction length”  $\lambda$  – the distance over which the initial number of hadrons reduces by a factor of  $1/e$ 
  - ▶ for high  $Z$  materials,  $\lambda$  is much greater than  $X_0$

material	$X_0$ [g/cm <sup>2</sup> ]	$\lambda$ [g/cm <sup>2</sup> ]
H <sub>2</sub>	63	52.4
Al	24	106
Fe	13.8	132
Pb	6.3	193

## Nuclear interaction length – remarks

- Nuclear interaction length is particle dependent – e.g.  $\lambda$  for  $\pi^+$  and  $\pi^-$  differ by  $\sim 5\%$
- the total energy deposited in a hadronic shower is not equal to the incident particle energy – a lot of it is lost due to escaping neutrinos
- hadron showers allow to capture neutrons which don't have electric charge and don't lose energy by ionization

# Range

- If we integrate the Bethe formula we'll get the mean distance travelled by a particle before it stops

- ▶ this number is called range  $R$

$$R = \int_0^R dx = - \int_{E_0}^0 \frac{dx}{dE/dx}$$

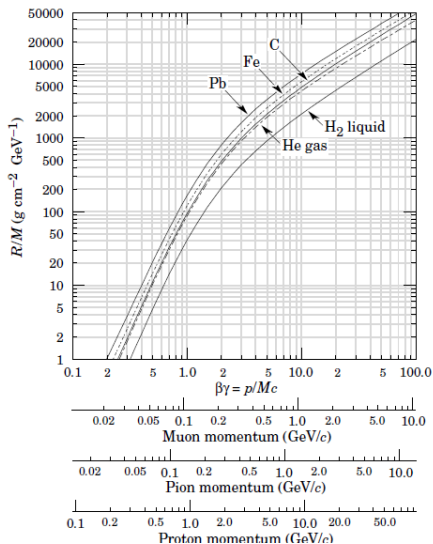
- ▶ Range is a useful concept only for low energy hadrons and muons

- For non-relativistic particles,  $-\frac{dE}{dx} = C \frac{Z z^2}{A v^2}$  where  $C$  is a constant (here I assume that  $\log v$  is a slowly changing function compared to  $1/v^2$  and so it is included in the constant) and  $E = m + T$ ,  $T = \frac{1}{2}mv^2$

$$R = \frac{A}{CZmz^2} \int_0^{T_0} 2T dT = \frac{AT_0^2}{CZmz^2} = \frac{Amv_0^4}{4CZz^2}$$

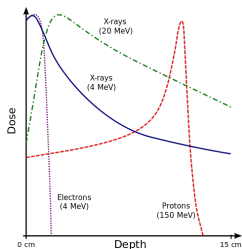
- Range is usually given in terms of  $R/m$  (independent from particle mass  $m$ )

# Range of heavy charged particles in various materials



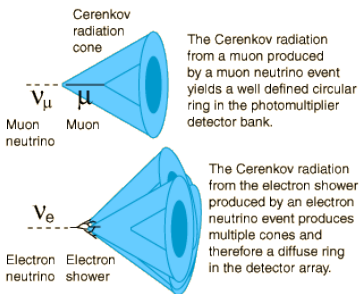
# Bragg curve

- As a heavy charged particle passes through matter, its energy decreases and the energy loss increases
  - ▶ most of the particle energy is deposited near the range  $R$ , leading to a sharp peak
- If energy losses were always equal to  $|dE/dx|$  as described by the Bethe formula, the particles would always travel a distance  $R$  and then stop
  - ▶ in practice, the deposited energy fluctuates, so the peak gets smeared
  - ▶ Landau distribution is bound on the left side, so the deposition peak does not extend much to the right
- The energy loss per unit length as a function of penetration depth is called the Bragg curve
  - ▶ knowing its exact shape is very important in particle therapy
- For electrons and photons, the energy losses are “catastrophic,” they travel a certain distance and then get kicked out
  - ▶ the energy deposited by  $e/\gamma$  decreases as an exponent of the distance, there is no peak



# Cherenkov radiation

- Cherenkov radiation is emitted when a charged particle goes with a speed larger than the speed of light in the medium
  - ▶ there is nothing that prevents a particle to do it as long as its speed is below  $c$
- The effect is similar to a sonic boom created by a super sonic aircraft moving faster than the speed of sound
  - ▶ actually this is not exactly true: a sonic boom is a non-linear effect, Cherenkov radiation is linear
- The radiation is concentrated in a cone of a half angle  $\theta = \arccos \frac{1}{n\beta}$ , where  $n$  is the refractive index
  - ▶ by measuring the size of the cone, it's possible to calculate the speed of the particle
- Particles in a electromagnetic cascade can produce what is called "coherent Cherenkov radiation"



# Transition radiation

- When a particle with charge  $ze$  goes from vacuum to a medium with plasma frequency  $\omega_p$ , it radiates energy  $\Delta E = \alpha z^2 \gamma \hbar \omega_p / 3$

$$\hbar \omega_p = \sqrt{\rho [\text{g/cm}^3] (Z/A)} \times 28.81 \text{ eV}$$

- ▶ Cherenkov radiation occurs when a particle goes in a medium with constant  $n$ ; transition radiation requires a change of  $n$
- Transition radiation can be used to determine particle's  $\gamma$ 
  - ▶ A popular use is to discriminate between electrons and heavy particles
- The amount of energy deposited in a single transition is very low, so for practical purposes one has to set up a stack of radiators separated by gas gaps